

# Irradiance Prediction Intervals for PV Stochastic Generation in Microgrid Applications

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## Abstract

The increasing interest in integrating volatile resources into microgrids implies the necessity of quantifying the uncertainty of photovoltaic (PV) production using dedicated probabilistic forecast techniques. The work presents a novel method to construct ultra-short-term and short-term prediction intervals (PIs) for solar global horizontal irradiance (GHI). The model applies the k-means algorithm to cluster observations of the clear-sky index according to the value of selected data features. At each timestep, the features are compared with the actual conditions to identify the representative cluster. The lower and upper bounds of the PI are calculated as the quantiles of the irradiance instances belonging to the selected cluster at a target confidence level. The validation is performed in 3 datasets of GHI measurements, each one of 85 days. The model is able to deliver high performance PIs for forecast horizons ranging from sub-second to intra-hour ahead without the need of additional sensing systems such as all-sky cameras.

*Keywords:* Solar forecasting, Prediction interval, Ultra-short term, k-means algorithm.

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## 1. Introduction

2 The thrust toward increasing the penetration of non-dispatchable renewable  
3 generation in the electrical grid requires to redefine conventional practices to

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4 assure reliable power system operation, see for example [1, 2]. A paradigm in-  
5 creasingly advocated in the recent technical literature to cope with the variabil-  
6 ity of stochastic generation is the development of robust and predictive controls.  
7 They take advantage of short-term forecasts of renewable generation in order  
8 to plan adequate counteractions to prevent, or mitigate, operational issues re-  
9 lated to renewables power fluctuations. Examples include the dispatchability  
10 of renewables, achieving self-consumption of locally generated electricity, and  
11 the short-term redispatching of conventional generation units, see for example  
12 [3, 4, 5, 6]. The period of the redispatch control action normally depends on  
13 the availability of the reserve in a given power grid and on the performance of  
14 the forecasting tools accounting for the uncertainties. For the case of micro-  
15 grids, their limited physical extension and the low granularity of the resources  
16 involve the necessity of coupling the reserves dispatch with their real-time con-  
17 trol. In this respect, a new protocol for real-time control of microgrids has  
18 been presented in the recent literature; in this framework, the control decision  
19 is updated with a sub-second resolution, [7, 8]. Since photovoltaic (PV) systems  
20 represent one of the major resources in modern microgrids, the availability of  
21 irradiance forecasting is beneficial to address the aforementioned challenges at  
22 forecast horizons from sub-second up to intra-hour, [2]. A further concern as-  
23 sociated with the dense penetration of PV installations in distribution systems  
24 and microgrids is the lack of the spatial smoothing effect, resulting in large vari-  
25 ations of the solar irradiance. As an example, Figures 1a and 1b respectively  
26 show daytime global horizontal irradiance measurements (GHI, recorded at the  
27 EPFL campus by using a pyranometer) and the power consumption of a group  
28 of EPFL buildings equipped with a 95 kWp PV-roof system. As visible, GHI  
29 variations (which varies up to 85% in magnitude in less than two minutes) cause  
30 very steep fluctuations of the power production and consumption. The avail-  
31 ability of high-quality ultra-short-term and short-term GHI forecast enables the  
32 possibility of taking preemptive control actions and mitigating the effect of its  
33 fluctuations.

34 In general, the choice of the forecast method is strictly related to the target

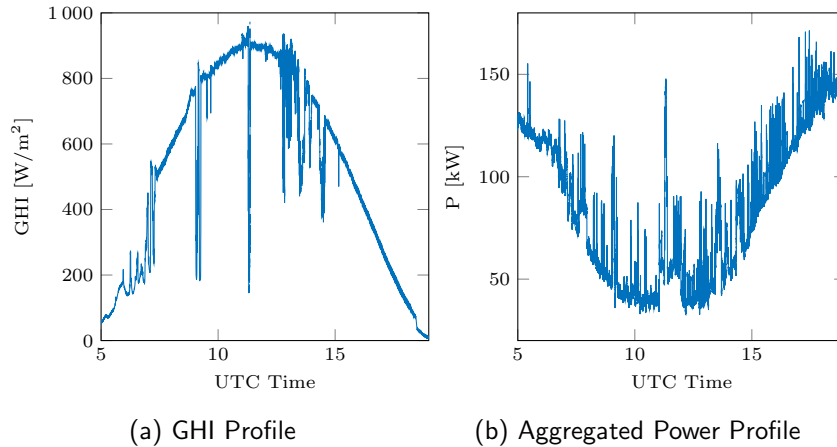


Figure 1: GHI and aggregated power profile (load and PV production) as a function of the UTC Time, registered at the EPFL campus on the 15<sup>th</sup> of May 2016.

35 forecast horizon and geographical scale. As explained in [9], day-ahead regional  
 36 irradiance forecasting relies on satellite observations and numerical weather pre-  
 37 dictions (NWPs). However, we here focus on local and shorter term forecasts  
 38 (lower than one hour) where Artificial Intelligence (AI) methods are generally  
 39 applied, [10]. The use of all-sky cameras is a promising solution for intra-hour  
 40 forecast horizons, as introduced in [11, 12]. To the best of our knowledge, the  
 41 only method addressing the problem of solar irradiance forecast at sub-second  
 42 time scale is the one proposed in [13].

43 Two main kinds of forecast are conceived: a first kind (deterministic) con-  
 44 siders only the point forecast while the second one (probabilistic) includes infor-  
 45 mation accounting for the intrinsic uncertainty of the prediction and it is more  
 46 appropriate when dealing with control and decision making in modern power  
 47 systems, [14]. Especially in the case of fast irradiance ramps, generally difficult  
 48 to predict, PIs are necessary to define the worst-case scenario that should be  
 49 considered in the control decision process.

50 Regarding GHI point predictions, the simplest forecast model is the persis-  
 51 tent one, which is commonly used as a benchmark for performance evaluation.  
 52 It assumes that the GHI remains constant with the forecast horizon. In gen-

53 eral, most of the point forecast techniques are based on AI methods. A more  
54 deterministic approach consists in detecting the position of the clouds, deduc-  
55 ing clouds motion and calculating the time when a cloud covers the sun, e.g.  
56 [11, 15]. Apart from cloud detection and motion, sky images contain more in-  
57 formation impacting the GHI prediction: examples are the cloud cover and the  
58 type of clouds. This kind of information can be combined with machine learning  
59 methods to compute the forecast, e.g. [16, 17].

60 Several works address the problem of probabilistic forecast and propose PIs  
61 computation models. Probabilistic solar power forecast is proposed in [18, 19]  
62 where a set of likely predictions (i.e. an ensemble) is provided using a historical  
63 set of variables and deterministic meteorological models. Authors of [18] use  
64 a distance criterion to retrieve similar past forecasts, under the assumptions  
65 that their errors are likely to be similar to the errors of the current forecast.  
66 These methods refer to 0-72 hours forecast horizons, considering hourly power  
67 data. In [20], a hybrid model is proposed, integrating Support Vector Machine  
68 (SVM), ANN and sky imaging techniques to deliver real-time PIs for direct  
69 normal irradiance (DNI) for 5, 10, 15, 20 minutes ahead. At each time step,  
70 the computational time is less than 5 seconds. Another stochastic approach in  
71 [21] proposes the design of a k-nearest neighbors (KNN) algorithm. The KNN  
72 algorithm is used to predict the GHI and DNI and their uncertainty intervals,  
73 for time horizons from 5 to 30 minutes. More recently, Authors of [22] proposed  
74 a data-driven method to construct GHI probability densities for one hour-ahead  
75 predictions, using nonparametric bootstrap and a map of solar position. The  
76 developed method has low computational complexity, requiring 0.56 seconds  
77 on a personal computer. In [23], point forecasts are generated using AutoRe-  
78 gressive Moving Average (ARIMA) and the associated PI is calculated using a  
79 Generalized AutoRegressive Conditional Heteroskedasticity model (GARCH),  
80 considering a prediction horizon from 10 minutes to 6 hours. The use of recur-  
81 sive formulas, to update the model parameters in real-time, allows to reduce the  
82 computational complexity of the method.

83 Having stated this, we note that the available literature lacks of a unique

84 forecasting tool for prediction horizons ranging from sub-second up to intra-  
85 hour and capable of operating at low levels of aggregation, where the level of  
86 volatility is higher due to the reduced spatial smoothing effect. While many  
87 methods have been proposed for intra-hour GHI forecasting and might be ap-  
88 plied to deliver ultra-short-term predictions, there is at least the compelling need  
89 of re-assessing their performance in the light of the requirements of real-time  
90 control of local power systems. Moreover, computational complexity becomes a  
91 key concern when considering the high reporting rate of ultra-short predictions,  
92 implying that available forecasting methods with complex on-line training pro-  
93 cedures (like ANNs, heuristic optimization-based and sky imaging) might not  
94 be suitable.

95 We propose a novel nonparametric method for ultra-short term forecasting  
96 of the global horizontal irradiance (GHI) to deliver predictions with a forecast  
97 horizon in the range from 500 ms to 5 minutes, thus suitable both in the context  
98 of real-time control of microgrids and energy management strategies. PIs com-  
99 putation is based on a well-known pattern recognition technique called k-means  
100 clustering, [24]. A training dataset is first clustered considering two empiri-  
101 cally selected influential variables. Then, PIs are calculated by extracting the  
102 quantiles of the cluster which resembles at most the actual conditions. A clear-  
103 sky model is also introduced for the de-trending of the GHI time series. The  
104 method does not require any information from sky-imaging since it only relies  
105 on measurements of the GHI, it is computationally efficient and needs a limited  
106 training dataset. As later qualified in the paper, the real-time generation of  
107 the PIs takes, for one time instance, less than 0.5 ms. Thus, the method is  
108 applicable even when the control decision has to be taken at sub-second time  
109 scale.

110 The paper is organized as follows: Section 2 defines the problem and intro-  
111 duces the nomenclature, Section 3 explains in details the methodology adopted  
112 to deliver the PIs and discusses its computational complexity. Section 4 de-  
113 scribes the available datasets and their characterization. Section 5 presents the  
114 results, comparing them with available benchmark methods. Section 6 shows

115 the main conclusions.

## 116 **2. Preliminaries on the Adopted Nomenclature**

117 PIs give a range of possible values in which the future realization is expected  
118 to lie with a given confidence level  $\alpha$ , [25]. At the timestep  $t$ , we denote the  
119 one-step-ahead PI at confidence level  $\alpha$  as composed by the upper and lower  
120 bounds:

$$\left( I_{t+1|t}^{\uparrow\alpha}, I_{t+1|t}^{\downarrow\alpha} \right). \quad (1)$$

121 Note that we do not have any assumption on the distribution of the time series  
122 since we adopt a nonparametric approach.

123 GHI measurements are pre-processed in order to remove the daily and sea-  
124 sonal components due to changes of the sun position. This is achieved by intro-  
125 ducing the clear-sky index  $K$ , which is defined as the ratio between the measured  
126 GHI and the clear-sky irradiance, respectively denoted by  $I$  and  $I_{cs}$ :

$$K = \frac{I}{I_{cs}}. \quad (2)$$

127 The clear-sky irradiance is the irradiance that would reach the ground in  
128 clear-sky conditions, i.e. absence of clouds. It is obtained by applying the clear-  
129 sky model implemented in the geographical information system GRASS, which  
130 also account for topological shading [26, 27].

## 131 **3. Methods**

132 The principle behind the proposed forecasting approach is that, if a real-  
133 ization of the solar irradiance happened in the past under certain measurable  
134 conditions, then it is likely to happen again under the same confluence of cir-  
135 cumstances.

136 The proposed PI estimation method consists in clustering historical data  
 137 according to the value of certain influential variables, introduced in the following.  
 138 The clusters are therefore used as empirical conditional probabilities of future  
 139 realizations and used to compute the PI by calculating the quantiles according  
 140 to a given confidence level. In particular, these influential variables should  
 141 be representative of the irradiance fluctuations since it is the main cause of  
 142 the uncertainty associated with solar forecasts. These variables, inputs of the  
 143 clustering process, are selected according to the literature that considers the  
 144 average and the variability of the clear sky-index as the most influential ones,  
 145 [21, 28]. We consider a training dataset of historical clear-sky index observations  
 146  $K_1, \dots, K_N$ , from which we extract the following influential variables:

- 147 • the average clear-sky index value on a mobile window of length  $n$  consid-  
 148 ering the most recent data points:

$$M_i = \frac{1}{n} \sum_{j=i-n}^i K_j, \quad i = n + 1, \dots, N \quad (3)$$

149 of which we consider the normalized version  $M_i^*$ . Namely, we normalize  
 150 the sequence  $M_0, \dots, M_1$  to a length of 1; <sup>1</sup>

- 151 • the clear-sky index variability:

$$V_i = \sqrt{\frac{1}{n} \sum_{j=i-n}^i (K_j - K_{j-1})^2}, \quad i = n + 1, \dots, N \quad (4)$$

152 which is a measurement of GHI fluctuations. As for the previous case, we  
 153 consider the normalized version  $V_i^*$ .

154 Normalization of the influential variables is required to enable a fair compar-  
 155 ison between parameters with different scale. For each observation, the vector  
 156  $\mathbf{p}_i$  of influential variables is:

$$\mathbf{p}_i = (M_i^*, V_i^*), \quad i = n + 1, \dots, N. \quad (5)$$

---

<sup>1</sup> The normalized version of a vector  $\mathbf{X}$  is a vector  $\mathbf{X}^* = \mathbf{X}/|\mathbf{X}|$ , where  $|\mathbf{X}|$  is the norm of  $\mathbf{X}$ .

157 The process to compute PIs is performed in two ways:

- 158 • Method A: we cluster the original clear-sky index time series;
- 159 • Method B: we cluster the differentiated clear-sky index time series:

$$\Delta K_i = K_i - K_{i-1}, \quad i = 2, \dots, N, \quad (6)$$

160 to verify if differencing leads to better prediction performance.

### 161 3.1. Clustering of the training set

162 The k-means iterative algorithm is firstly used to classify historical observa-  
163 tions of clear-sky index according to predefined influential variables. K-means  
164 clustering is a partitioning algorithm that allocates each observation into ex-  
165 actly one of the  $k$  clusters, each one defined by a representative centroid. In  
166 particular,  $k$  centroids are at first randomly selected (the first centroids are sim-  
167 ply uniformly random observations). Then, each vector of the training dataset  
168 is assigned to the closest centroid, and the centroid is iteratively recalculated as  
169 the mean of the vectors of each class until convergence is reached (i.e., centroids  
170 do not change anymore between iterations).

#### 171 *Method A*

172 We apply the k-means algorithm to cluster the vectors  $\mathbf{p}_i$  belonging to the  
173 training set, being  $k$  the number of clusters. The algorithm assigns to each vector  
174  $\mathbf{p}_i$  a cluster index  $l$  between 1 and  $k$  and determines the centroids locations  
175  $\mathbf{c}_l = (M_l^*, V_l^*)$  for  $l = 1, \dots, k$ .

176 We denote the generic cluster  $G_l$  as composed by all the clear-sky indexes  
177  $K_{i+1}$  for which  $\mathbf{p}_i$  has index  $l$ .

#### 178 *Method B*

179 We apply the same clustering procedure described above for Method A.

180 However, we denote the generic cluster  $\Delta G_l$  as composed by all the differ-  
181 entiated clear-sky realizations  $\Delta K_{i+1}$  for which  $\mathbf{p}_i$  has index  $l$ .

182



183 An example of the k-mean partitioning of the influential variables is shown  
 184 in Fig. 2 where the normalized clear-sky index average and variability are  
 185 clustered.

186 It is worth noting that the k-means clustering of the training dataset can  
 187 be performed off-line on historical data. This is a key aspect if considering the  
 188 high reporting rate of predictions for microgrid applications since it allows to  
 189 reduce the computational complexity.

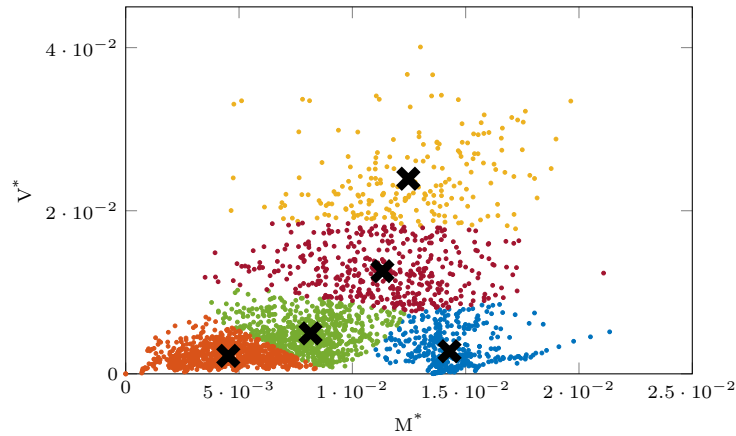


Figure 2: Example of k-means clustering obtained for  $k = 5$ . The  $x$  and  $y$  axis represent the normalized clear-sky index average and variability, respectively. The black marker signs the centroids.

### 190 3.2. Prediction Intervals

191 In this section we describe how PIs are computed, distinguishing between  
 192 the two proposed methods.

#### 193 Method A

194 Starting from the clusters  $G_1, \dots, G_k$  defined in the previous section, the  
 195 PIs at the target confidence level  $\alpha$  can be computed as:

$$q_l^{\uparrow\alpha} = (1 + \alpha)/2 \text{ quantile of } G_l, \quad l = 1, \dots, k \quad (7)$$

$$q_l^{\downarrow\alpha} = (1 - \alpha)/2 \text{ quantile of } G_l, \quad l = 1, \dots, k. \quad (8)$$

196 For increased computational efficiency, we note that also this operation can be  
 197 performed off-line, and the PIs for each class can be stored.

198 Say being at time  $t$ , the objective is to perform the on-line computation of  
 199 the PI for the next time interval  $t + 1$ . The vector of influential variables at  $t$   
 200 is denoted by  $\mathbf{p}_t = (M_t^*, V_t^*)$ . It is calculated normalizing the raw influential  
 201 variables  $M_t, V_t$  with respect to those available in the training data set. The next  
 202 step is the calculation of the Euclidean distances between  $\mathbf{p}_t$  and the centroids  
 203  $\mathbf{c}_l$ :

$$d_l = \|\mathbf{c}_l - \mathbf{p}_t\|^2, \quad l = 1, \dots, k \quad (9)$$

204 which is used as a similarity criterion to select the cluster representative of the  
 205 future clear-sky outcome. We indicate with  $\hat{l}$  the index corresponding to the  
 206 cluster with minimum distance. It is used to select the quantiles used in the PI  
 207 computation as:

$$K_{t+1|t}^{\uparrow\alpha} = q_l^{\uparrow\alpha}. \quad (10)$$

$$K_{t+1|t}^{\downarrow\alpha} = q_l^{\downarrow\alpha} \quad (11)$$

#### 208 *Method B*

209 Starting from the clusters  $\Delta G_1, \dots, \Delta G_k$  obtained from the differentiated  
 210 GHI time series, the PIs at the target confidence level  $\alpha$  can be computed as:

$$q_l^{\downarrow\alpha} = (1 - \alpha)/2 \text{ quantile of } \Delta G_l, \quad l = 1, \dots, k \quad (12)$$

$$q_l^{\uparrow\alpha} = (1 + \alpha)/2 \text{ quantile of } \Delta G_l, \quad l = 1, \dots, k. \quad (13)$$

211 Also in this case, the quantiles extraction is computed off-line. The on-line  
 212 computation of the PIs consists in finding the index  $\hat{l}$  of the cluster with centroid  
 213 at the minimum distance from  $\mathbf{p}_t$ . It is used to select the quantiles used in the  
 214 PI computation as:

$$K_{t+1|t}^{\uparrow\alpha} = K_t + q_l^{\uparrow\alpha}, \quad (14)$$

$$K_{t+1|t}^{\downarrow\alpha} = K_t + q_l^{\downarrow\alpha}. \quad (15)$$

215 i.e., the current measurement is summed to the upper and lower quantiles of  
216 the differentiated time series.

217

218 It is important to note that, so far, PIs are computed to forecast the clear-sky  
219 index. The last step consists in computing the PI for the GHI:

$$\hat{I}_{t+1|t}^{\uparrow\alpha} = K_{t+1|t}^{\uparrow\alpha} J_{cs,t+1}, \quad (16)$$

$$\hat{I}_{t+1|t}^{\downarrow\alpha} = K_{t+1|t}^{\downarrow\alpha} J_{cs,t+1}. \quad (17)$$

220 In the results section, performance of methods A and B are evaluated by com-  
221 paring the estimated PIs as defined in (16)-(17) with the GHI measurements.

### 222 3.3. Selection of the Parameters

223 The parameters we need to specify when applying the k-means clustering  
224 procedure are:

- 225 • The number of samples  $n$  used in (3) and (4);
- 226 • The number of cluster  $k$  used for the partition of the training dataset;
- 227 • The length of the training dataset  $N$ .

228 The selection of the parameters values is a sensitivity process that is exhaustively  
229 evaluated in Section 5. The assessment is performed in a searching dataset,  
230 and then the selected values are applied in a testing dataset for performance  
231 evaluation.

232 While it was seen from the results that variations of  $n$  in the range from 2  
233 to 5 do not, in general, alter modeling performance, the values of  $k$  and  $N$  are  
234 interdependent and the selection of these two parameters needs to be carried out  
235 simultaneously (i.e. we need to find the combination of  $k$  and  $N$  with the best  
236 performance). For each specific case, a diagnostic analysis should be performed.  
237 We consider here two main approaches to fix  $k$  and  $N$  as described in the next  
238 subsections.

239 *3.3.1. A-posteriori Selection of  $k$  with Exhaustive Search (ES)*

240 Several  $(k, N)$  combinations are attempted in a searching dataset: the can-  
241 didate combination is the one with best a-posteriori prediction performance.  
242 These values are then used to evaluate the performance of the testing dataset,  
243 supposing that it exhibits similar characteristics of the searching set. This ap-  
244 proach can become computationally expensive, therefore motivating the devel-  
245 opment of methods for the a-priori selection of the free parameters, as described  
246 in the next subsection.

247 *3.3.2. A-priori Selection of  $k$  with Silhouette Analysis (SA)*

248 The objective is to use Silhouette Analysis, [29], to improve the partitioning  
249 of the training dataset, allowing for an a-priori selection of parameter  $k$  and  
250 in order to avoid the exhaustive approach. The Silhouette Analysis consists in  
251 some main steps:

- 252 • a small value of  $k$  is chosen (e.g.  $k = 5$ ) and the clustering algorithm is  
253 run;
- 254 • the silhouette value for a generic point  $i$  is calculated as:

$$s(i) = \frac{a(i) - b(i)}{\max(a(i), b(i))} \quad (18)$$

255 where  $a(i)$  is the average distance from point  $i$  to the other points in the  
256 same cluster, while  $b(i)$  is the minimum average distance from instance  $i$   
257 to points in a different cluster, minimized over clusters. Parameter  $s$  is a  
258 measure of how close the instance is to the other instances in its cluster  
259 and how far it is to those in the other clusters. In general, a silhouette  
260 value close to 1 is desired because it means that the point is well clustered  
261 while a value close to -1 means misclassification;

- 262 • the mean of the silhouette values is computed. If most points have a high  
263 silhouette value, then the clustering is appropriate;
- 264 • the value of  $k$  is augmented and it is evaluated if having more clusters  
265 allows for a better partitioning (higher mean of the silhouette values);

266 • the number of clusters is selected equal to the value  $k$  above which we do  
267 not see any improvement in terms of increasing of the average silhouette  
268 value.

269 As shown in subsection [5.3.2](#) the value of  $N$  required to converge at constant  
270  $k$  is not sensitive to the characteristics of the dataset. In general, for each  
271 forecast horizon and given  $k$ , it is possible to identify a value of  $N$  above which  
272 performance is close to convergence. Above this value, small oscillations are  
273 explained by the intrinsic stochasticity of the data. This feature is important  
274 for the modeler since it allows for the reduction of the parameters to be found.

### 275 *3.4. Algorithms time complexity*

276 In this section, we evaluate the time complexity of the real-time computation  
277 of the proposed algorithms. This is an important aspect because they are de-  
278 signed with the stated objective of delivering PIs to real-time control processes  
279 for electrical power systems.

280 The algorithms consist of two parts, the training phase and on-line compu-  
281 tation of PIs. The former does not have any real-time requirement and can be  
282 performed off-line. The latter phase is instead time critical, and it is to per-  
283 form with a hard-real deadline. First, it consists in calculating the normalized  
284 influential variables, [\(3\)](#)-[\(4\)](#), an operation with constant time complexity,  $O(1)$ ,  
285 which involves algebraic operations. Then, we have the computation of  $k$  norms,  
286  $O(k)$ , and a minimum search, which can be performed efficiently with a merge  
287 search,  $O(k \log(k))$ . Considering that  $k$  is fixed by design, the time complexity  
288 of the real-time computation is constant time,  $O(1)$ , therefore denoting that  
289 complexity does not grow with the size of the problem (scalability). Statistics  
290 on the execution time of the algorithms are given in Section [5.8](#).

## 291 **4. Data**

### 292 *4.1. Data acquisition*

293 Global horizontal irradiance measurements are collected at 50 ms resolu-  
294 tion at the École Polytechnique Fédérale de Lausanne (EPFL) by using an

295 Apogee SP-230 all-seasons pyranometer which is located at the following GPS  
296 coordinates: 46.518397-N, 6.565229-E. We consider three datasets of 85 days  
297 each, corresponding to different periods of the year. The first contains irradi-  
298 ance measurements from July to September 2015 (Summer), the second from  
299 October to December 2015 (Autumn), and the third from January to March  
300 2016 (Winter). Each dataset is divided into a searching subset of 55 days and  
301 a testing one composed of the remaining 30 days.

302 The original time series is down-sampled to three different resolutions: 500  
303 ms, 1 minute and 5 minutes. These series are used to compute one-step-ahead  
304 PIs for the corresponding forecast horizon. Down-sampling is computed by  
305 averaging the intermediate samples.

306 It is worth noting that applying the clear-sky normalization causes very high  
307 values of  $K$  close to sunrise and sunset. Therefore, we consider only daylight  
308 values covering the period of the day for which the clear-sky index does not  
309 diverge.

#### 310 4.2. Data classification

311 Characterizing the dataset is important for performance comparison and  
312 evaluation. Indeed, the robustness of the method should be tested during pe-  
313 riods of different irradiance volatility. In our case, we are interested in charac-  
314 terizing the three available datasets: Summer, Autumn, and Winter. First, we  
315 give an information regarding the weather of the selected period and location.  
316 In particular, we retrieve cloud cover data <sup>[2]</sup> from MeteoSwiss Idaweb services,  
317 <sup>[30]</sup>, from two weather stations in the vicinity of our installation. The average  
318 cloud okta values for the three seasons are: 3.86 okta in Summer, 4.67 okta in  
319 Autumn and 5.96 okta in Winter.

320 To be more specific, we introduce a criterion consisting in counting the per-

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<sup>2</sup>Cloud cover corresponds to the fraction of the sky obscured by clouds when observed from a given location. The unit of measurement is the okta, ranging from 0 (completely clear-sky) to 8 (completely overcast).

321 centage of periods with a volatility lower than a given threshold. For each  
 322 timestep  $t$ , we calculate the per-unit difference as  $\Delta I_t = (I_t - I_{t-1})/I_{max}$ , where  
 323  $I_{max} = 1000 \text{ W/m}^2$ . For each prediction horizon, we establish a threshold for  
 324  $\Delta I_t$ , above which the observation at time  $t$  is considered with high volatility.  
 325 The threshold is empirically computed as the 99% quantile of the  $\Delta I$  time se-  
 326 ries obtained by manually selecting a period of 3 clear-sky days. The values are  
 327 shown in Table 1 for different forecast horizons.

Table 1: Thresholds.

Time Horizon	Thresholds
500 ms	0.0004
1 min	0.011
5 min	0.025

Table 2: Percentage of periods with high irradiance volatility.

Season	Forecast Horizon		
	500 ms	1 min	5 min
Summer	16	17	22
Autumn	5	9	12
Winter	13	15	20

328 The percentage of periods exceeding the threshold of GHI high volatility is  
 329 shown in Table 2 for each dataset and for different forecast horizons. As it can  
 330 be observed, the Summer period is characterized by the highest GHI volatility,  
 331 followed by Winter and Autumn.

## 332 5. Results and Discussion

333 First, the metrics used for performance evaluation are introduced in Subsec-  
334 tion 5.1. Then, Subsection 5.2 shows the advantage given by the introduction  
335 of a clear-sky model at different forecast horizons. In Subsection 5.3 the sensi-  
336 tivity of the performance with respect to the selection of the model parameters  
337 is discussed. In Subsections 5.4-5.6 the performance of the proposed methods  
338 is benchmarked against existing techniques. First, we compare the proposed  
339 methodology with the symmetric quantile extraction, which is the simplest way  
340 to construct our intervals. We use the empirical quantiles extracted from the  
341 distribution of the time series to build the PIs as in (10)-(11) and (14)-(15), re-  
342 spectively. It is important to highlight that the quantiles at time  $t$  are extracted  
343 from the whole time series, from  $t = 0$  to  $t - 1$ .

344 As a second benchmark, we compare our method with a model commonly used  
345 in forecasting. We first generate a point forecast using AutoRegressive Moving  
346 Average model (ARIMA), [31], with Double Exponential Smoothing. Then PIs  
347 are constructed assuming a Gaussian distribution of the point forecast error as:

$$K_{t+1|t}^{\uparrow\alpha} = \hat{K}_{t+1} + \eta_{\alpha}\sqrt{\sigma_t}, \quad (19)$$

$$K_{t+1|t}^{\downarrow\alpha} = \hat{K}_{t+1} - \eta_{\alpha}\sqrt{\sigma_t}. \quad (20)$$

348 where  $\hat{K}_{t+1}$  is the point forecast obtained by the ARIMA model,  $\eta_{\alpha}$  is the  
349 quantile of the normal distribution corresponding to the target confidence level  
350  $\alpha$  and  $\sigma_t$  is the variance of the forecast error.

351 Unless otherwise indicated, the target confidence level used for the following  
352 analysis is fixed equal to 95%.

353 Subsection 5.7 presents and discusses the reliability diagrams. Finally, statis-  
354 tics of the method execution time are provided in Subsection 5.8.

### 355 5.1. Metrics

356 We use three standardized metrics from the existing literature to evaluate the  
357 performance of the proposed methods [20, 32]. The first metric is the PI coverage



358 probability (PICP) which counts the number of times that the realization falls  
 359 inside the PI for a given confidence level  $\alpha$ :

$$\text{PICP} = \frac{1}{L} \sum_{t=1}^L c_t \quad (21)$$

360 where  $L$  is the total number of forecast instances of the testing dataset and

$$c_t = \begin{cases} 1, & \hat{I}_{t+1|t}^{\downarrow\alpha} \leq I_{t+1} \leq \hat{I}_{t+1|t}^{\uparrow\alpha} \\ 0, & \text{otherwise.} \end{cases} \quad (22)$$

361 Then, to account for the fact that the wider the PI, the easier it is to have  
 362 a realization falling inside it, we measure the PI normalized averaged width  
 363 (PINAW):

$$\text{PINAW} = \frac{1}{L I_{max}} \sum_{t=1}^L (\hat{I}_{t+1|t}^{\uparrow\alpha} - \hat{I}_{t+1|t}^{\downarrow\alpha}), \quad (23)$$

364 where  $I_{max} = 1000 \text{ W/m}^2$ . The value has been selected accordingly to the work  
 365 in [21], that is later used as benchmark for performance comparison. The third  
 366 metric quantifies the trade-off between having a large coverage probability and  
 367 small interval width. It is called coverage width-based criterion (CWC):

$$\text{CWC} = \text{PINAW}(1 + \gamma(\text{PICP})e^{-\mu((\text{PICP})-\mu)}) \quad (24)$$

368 where

$$\gamma = \begin{cases} 0, & \text{PICP} \geq \alpha \\ 1, & \text{PICP} < \alpha. \end{cases} \quad (25)$$

369 The parameter  $\mu$  can be tuned based on how much bad PIs are to penalize,  
 370 see [32]. We select here  $\mu = 10$ . PIs should have high PICP (higher or equal to  
 371 the target confidence level) coupled with a low value of PINAW.

## 372 5.2. Clear-sky index and GHI time-series comparison

373 A first analysis aims at assessing the difference between using the measured  
 374 irradiance (GHI) or the clear-sky index ( $K$ ) time series as inputs for the PIs

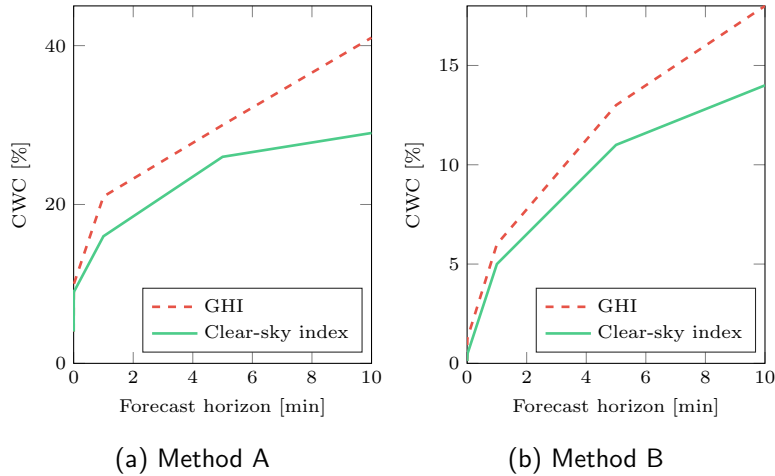


Figure 3: CWC as a function of the time horizon. Comparison between the use of the GHI time series (original) and the clear-sky index one (normalized).

375 computation method. In particular, we apply methods A and B and we increase  
 376 the forecast horizon to evaluate when the inclusion of a clear-sky model becomes  
 377 advantageous.

378 Fig. 3 shows that the CWC is, in general, lower (better performance) when  
 379 using the clear-sky index. In particular, the advantage of using a clear-sky  
 380 model becomes evident for time horizons longer than 1 minute. As expected,  
 381 the normalization of the time series becomes more important at higher fore-  
 382 cast horizons, when the effect of the sun position becomes more dominant.  
 383 When referring to ultra-short term horizons, fluctuations of solar irradiance are  
 384 mainly related to cloud motion and the importance of a clear-sky model becomes  
 385 marginal. Since the inclusion of a clear-sky model only leads to similar or bet-  
 386 ter prediction performance, the proposed methods are applied to the clear-sky  
 387 index time series.

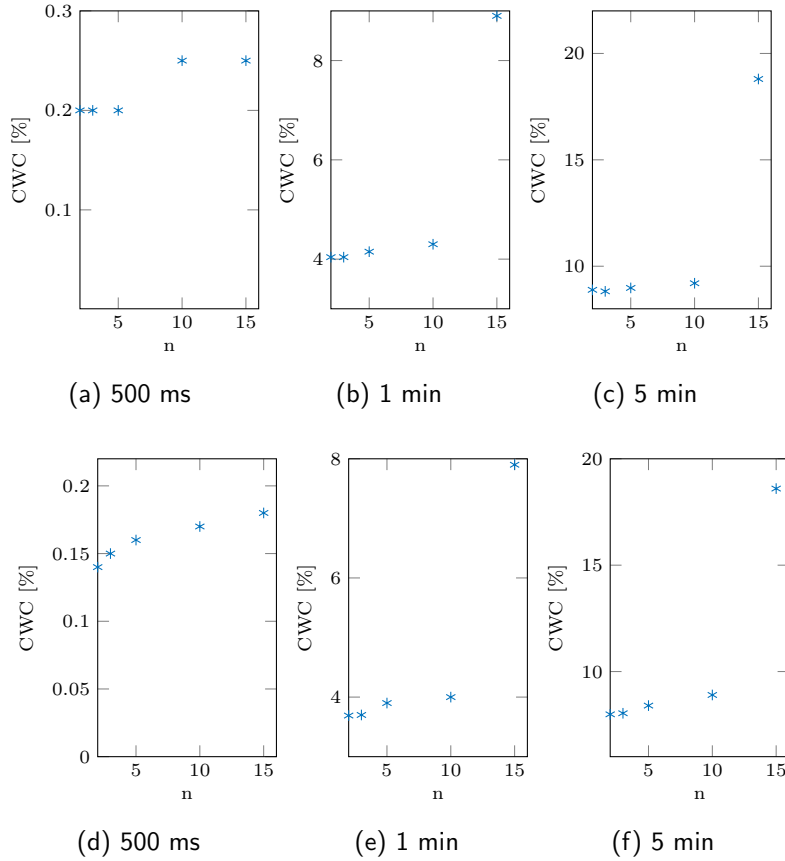


Figure 4: CWC as a function of  $n$  for different forecast horizons.  $N$  is equal to 30 days.  $k$  is equal to 5 for cases (a)-(c) and equal to 30 for case (d)-(f).

### 388 5.3. Parameters Selection and Sensitivity analysis

#### 389 5.3.1. Selection of parameter $n$

390 Fig. 4 shows the CWC metric as a function of  $n$  (3-4) obtained from a-  
 391 posteriori analysis of the performance of the whole Autumn dataset. Method B  
 392 is applied. The analysis considers different forecast horizons for both  $k=5$ , cases  
 393 (a)-(c), and  $k=30$ , cases (d)-(f). It is possible to see that performance is not  
 394 very sensitive to variations of  $n$  in the range from 2 to 5. From an a-posteriori  
 395 analysis of our datasets at different forecast horizons, we can conclude that  $n$   
 396 can be fixed to a value between 2 and 5, for all the considered cases. Indeed,

397 analogous conclusions can be inferred for Method A and different datasets (not  
398 shown here because of the similar behavior). These values of  $n$  are a good trade-  
399 off between having enough significant measurements to compute the influential  
400 variables and avoiding to consider realizations that are too far from the actual  
401 conditions. The results presented in what follows are obtained with  $n = 3$ .

### 402 5.3.2. Selection of parameter $k$ and $N$

403 As explained in Subsection 3.3 two main procedures are proposed to assign  
404 parameters  $k$  and  $N$ . In the case of the exhaustive search, we isolate 55 days  
405 of each dataset to perform the exhaustive searching. The number of clusters  $k$   
406 can vary between one (single cluster) and the total number of training samples  
407 (each data is assigned to its own cluster). Since the computational effort of the  
408 k-means algorithm is linearly dependent to the number of clusters and to the  
409 number of data, we here limit  $k$  to 1000 and  $N$  to 30 days. The dashed lines in  
410 Figures 5, 6 and 7 show the value of  $k$  that returns the best performance for a  
411 fixed  $N$ . This  $k$  is obtained a-posteriori by applying methods A and B to the  
412 searching dataset of 55 days ( $k_{ES,A}$  and  $k_{ES,B}$ , respectively). The three figures  
413 refer to different datasets. Each figure includes three sub-figures referring to  
414 the three forecast horizons, respectively. It is possible to note that, due to the  
415 heuristic nature of the methods, the optimal value of  $k$  cannot be known a-priori  
416 and it varies among the different considered cases (namely, we do not have a  
417 global optimum). Therefore, we select  $k$  and  $N$  as the combination returning  
418 the best prediction performance (minimum CWC), found a-posteriori. These  
419 values found for the searching set of 55 days are then applied for performance  
420 evaluation in the remaining 30 days.

421 In the case of the silhouette analysis,  $k$  is calculated for different  $N$  as the  
422 one maximizing the average silhouette of the training set and it is shown in  
423 Figures 5, 6 and 7 with the solid line. The value of  $k$  that maximizes the  
424 average silhouette does not correspond to the one returning the best forecasting  
425 performance. However, we notice that its value does not vary with  $N$  and can  
426 be selected independently. In order to determine  $N$  in the case of the silhouette

427 analysis, we fix  $k$  equal to the one returned by the analysis ( $k=5$ ) and we evaluate  
 428 the prediction performance for a different number of training days. Figures 8, 9  
 429 and 10 show the CWC (in logarithmic scale) as a function of  $N$ . They refer to  
 430 500 ms, 1 and 5 minutes, respectively. Each figure consists of two plots, showing  
 431 the performance for Method A and B, respectively. We can make the following  
 432 observations:

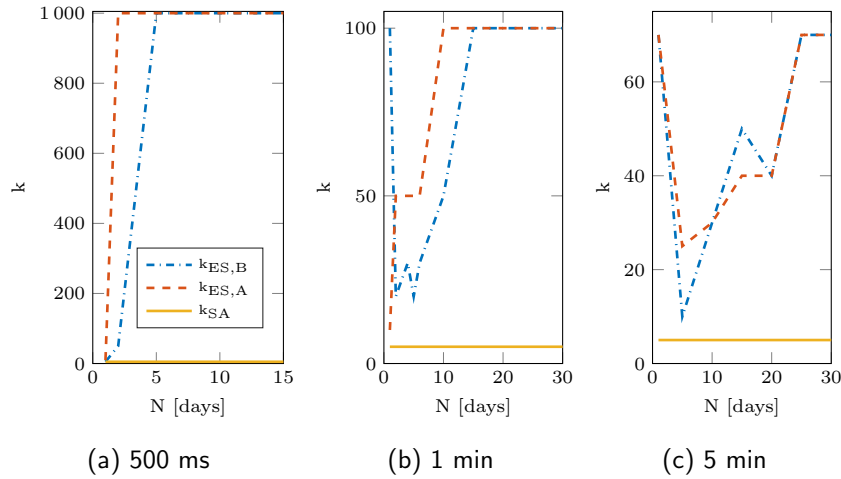


Figure 5: Number of clusters  $k$  as a function of the length of the training dataset. The Summer dataset is selected for the analysis. The dashed lines refer to the value of  $k$  corresponding to maximum performance for Method A and B, respectively. It is calculated a-posteriori by applying the ES. The solid line refers to the value of  $k$  from the SA.

- 433 • For each forecast horizon, it is possible to identify a first drop of CWC  
 434 after which performance tends to stabilize. The value of  $N$  that leads  
 435 to performance stabilization is not sensitive to the dataset and can be  
 436 fixed independently. On the contrary, as it is shown in the next sections,  
 437 performance at convergence depends on the nature of the dataset and, in  
 438 general, the behavior of the PIs depends on the volatility content of the  
 439 dataset.
- 440 • For the sub-second time horizon and Method A we have a first drop  
 441 of CWC after about one day and then performance tends to stabilize.

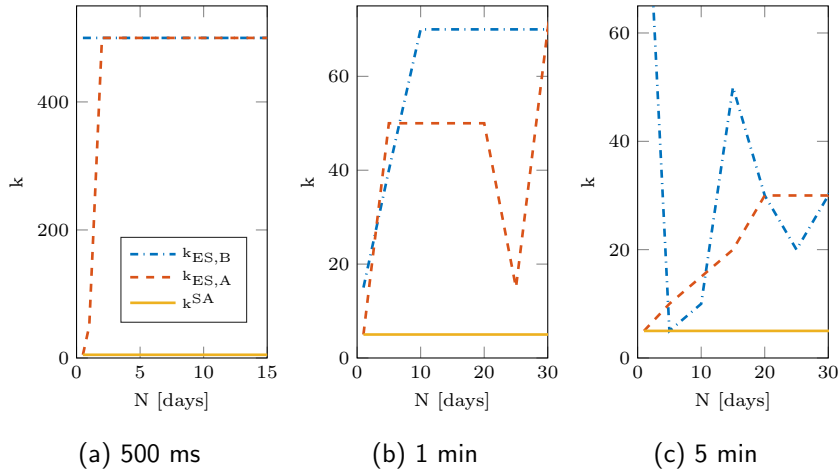


Figure 6: Number of cluster  $k$  as a function of the length of the training dataset. The Autumn dataset is selected for the analysis.

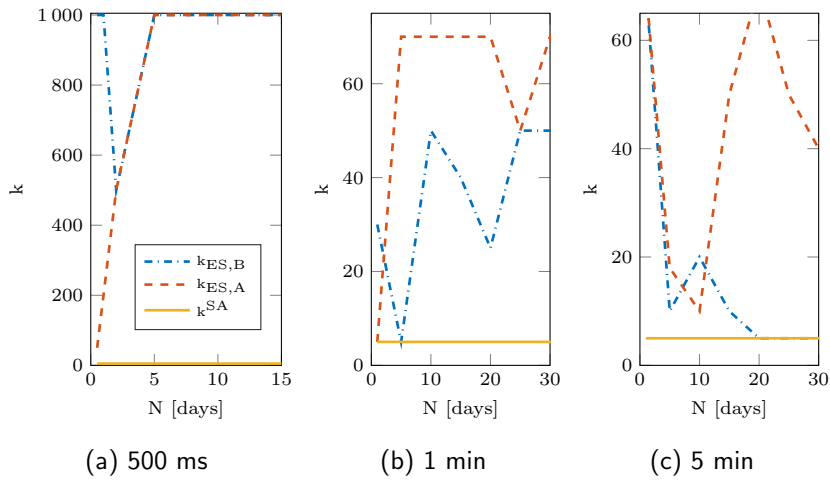


Figure 7: Number of cluster  $k$  as a function of the length of the training dataset. The Winter dataset is selected for the analysis.

442 Method B reaches convergence after few hours of training with subsequent  
 443 small CWC oscillations.

- 444 • For time horizon of 1 minute we have a first drop of CWC after about 5  
 445 days of training and then performance smooths out more slowly.

446 • For time horizon of 5 minutes we have a first drop of CWC after about 5  
447 days of training and then a second drop after 10 days. Then, performance  
448 smooths out more slowly.

449 In conclusion, when applying the exhaustive search we use the optimal  
450 combination of  $k$  and  $N$  found for the 55 days dataset as candidates for  
451 performance evaluation in the remaining 30 days. On the contrary, when  
452 applying the silhouette analysis approach, we select  $k=5$  and  $N$  equal to  
453 1, 5 and 10 days respectively for three time horizons. These values are  
454 valid for all the three datasets. A comparison between the two approaches  
455 is presented in what follows.

#### 456 5.4. Ultra-short term forecasting: Performance Evaluation

457 We focus here on sub-second forecast horizon: one-step-ahead PIs at 500 ms.  
458 At first, we analyze the two proposed approaches to compute  $k$  and  $N$ . Then,  
459 performance of methods A and B is evaluated, and therefore compared with  
460 existing methods from the technical literature.

##### 461 5.4.1. Exhaustive Search and Silhouette Analysis

462 In this section, we compare the performance obtained by applying the ex-  
463 haustive search and the silhouette analysis to determine  $k$  and  $N$ . Furthermore,  
464 results are compared with the optimal performance found a-posteriori to evalu-  
465 ate how far the estimations are from the optimum. Evaluation is carried out in  
466 the testing set of 30 days, for each one of three datasets.

467 Results are shown in Tables 3 for methods A and B. The comparison con-  
468 sider metric CWC. For 500 ms forecast horizon, the exhaustive search is out-  
469 performing the silhouette analysis and returns performance very close to the  
470 optimal a-posteriori. Indeed, when dealing with high sampling frequency, the  
471 large amount of data would require a number of clusters which is much higher  
472 than the one returned by the silhouette analysis. For sub-second time horizons,  
473 we proceed our analysis by using the exhaustive search method.

474 The following additional conclusions can be drawn from this analysis:

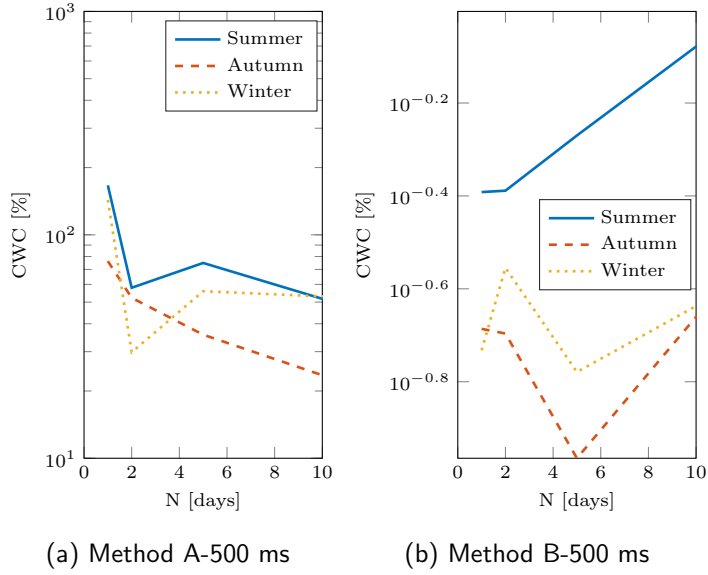


Figure 8: CWC as a function of the number of training days for 500 ms time horizon. CWC is shown in logarithmic scale.

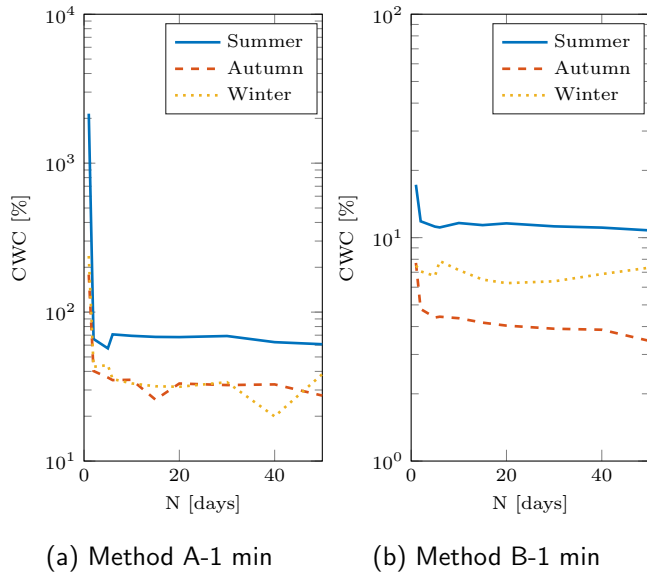


Figure 9: CWC as a function of the number of training days for 1 min time horizon. CWC is shown in logarithmic scale.



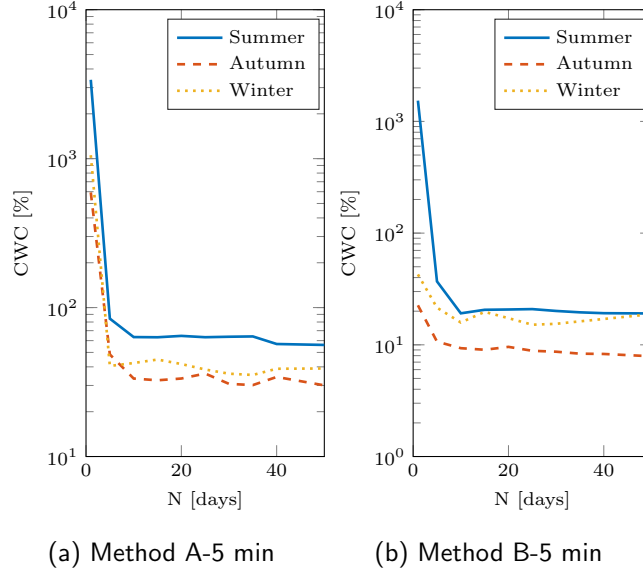


Figure 10: CWC as a function of the number of training days for 5 min time horizon. CWC is shown in logarithmic scale.

Table 3: CWC [%] for 500 ms,  $\alpha=95\%$ .

	Season				Season		
Method	Summer	Autumn	Winter	Method	Summer	Autumn	Winter
Optimal	2.88	1.18	2.55	Optimal	0.24	0.046	0.13
Silhouette	51	27.7	20.3	Silhouette	0.37	0.12	0.27
Exhaustive Search	4.69	2.41	4.03	Exhaustive Search	0.27	0.047	0.15

(a) Method A

(b) Method B

- 475 • For ultra-short term forecast, Method B outperforms Method A for each  
 476 considered case.
- 477 • For both the methods, the Summer period is characterized by worse per-  
 478 formance and this is explained by the highest volatility content, as shown  
 479 in Table 2. As expected, the Autumn period returns the best performance.

480 *5.4.2. Comparison with Benchmark Methods*

481 Table 4 shows the performance of the proposed methods compared with the  
482 above described benchmarks. First, we refer to the simple quantiles extraction  
483 (Quantiles A and Quantiles B), where the quantiles are computed from the  
484 original and differentiated time series respectively, as described in the introduc-  
485 tion of the results section. This comparison aims at showing the performance  
486 improving obtained by the k-mean clustering compared to the case where we  
487 extract the quantiles of the whole time series, without any clustering process.  
488 We can conclude that the k-means clustering is beneficial and leads to relevant  
489 performance improvement for all the analyzed cases.

490 The last row of Table 4 shows the results obtained by applying the ARIMA  
491 model and assuming a Gaussian distribution of the point forecast error (ARIMA  
492 + GAUSS). This method has to be compared with our proposed Method B since  
493 it requires a point forecast to compute the PIs. For 500 ms forecast horizon,  
494 the model is over confident with respect to the assumed normal distribution,  
495 returning PICP higher than 99% for  $\alpha=95\%$ . Thus, to allow a fair comparison  
496 with our method, we empirically adjust the target confidence level (and so  $\eta$ ) in  
497 order to obtain values of PICP similar to those given by the k-mean algorithm.  
498 Table 4 shows that, for the same coverage probability, Method B is characterized  
499 by lower PINAW.

500 We refer to [13] as the only reference method for ultra-short term available  
501 in the literature. To allow a fair comparison, we compare Method B with the  
502 Dynamic Interval Predictor (DIP) coupled with the persistent point forecast,  
503 proposed in [13]. Indeed, the DIP needs to be coupled with a point forecast  
504 technique.

505 From Table 5 we can conclude that the proposed method shows better per-  
506 formance when compared to the literature with respect to ultra-short term hori-  
507 zons.

Table 4: PICP-PINAW-CWC [%] for a time horizon of 500 ms,  $\alpha=95\%$ .

	Season		
Method	Summer	Autumn	Winter
Method A	90.5-1.94-4.69	93.7-0.33-2.41	92.6-1.85-4.03
Quantiles A	94.6-57.4-113	93.3-29.5-62.6	95.7-35.1-35.1
Method B	97.0-0.27-0.27	96.1-0.047-0.047	98.2-0.15-0.15
Quantiles B	90.4-0.35-0.35	91.4-0.13-0.30	91.0-0.12-0.28
ARIMA+GAUSS	97.0-0.50-0.50	96.1-0.1-0.1	98.2-0.32-0.32

Table 5: PICP-PINAW-CWC [%]. Performance comparison of the proposed Method B with the Dynamic Interval Predictor, [13].  $\alpha=95\%$ .

	Season		
Method	Summer	Autumn	Winter
Method B	97.0-0.27-0.27	96.1-0.047-0.047	98.2-0.15-0.15
DIP	97.2-0.36-0.36	96.0-0.053-0.053	97.4-0.19-0.19

508 *5.5. Short term forecasting*

509 In this section, we extend the proposed methods to higher forecast horizons  
 510 (i.e. minutes).

511 *5.5.1. Exhaustive Search and Silhouette Analysis*

512 Tables 6 and 7 show metric CWC obtained by applying the exhaust search-  
 513 ing, the silhouette analysis and the optimum a-posteriori, for 1 and 5 minutes  
 514 forecast horizons. The silhouette analysis coupled with Method B shows here  
 515 the best performance and is used for further comparison. On the contrary, when  
 516 using the original time series, the exhaustive search performs better and should  
 517 be used to select  $k$  and  $N$ .

Table 6: CWC [%] for 1 min,  $\alpha=95\%$ .

	Season				Season		
Method	Summer	Autumn	Winter	Method	Summer	Autumn	Winter
Optimal	21.7	8.50	22.9	Optimal	6.82	2.83	6.20
Silhouette	58.0	31.6	33.0	Silhouette	10.5	3.26	9.10
Exhaustive Search	25.2	11.4	23.4	Exhaustive Search	14.0	3.81	10.3

(a) Method A

(b) Method B

Table 7: CWC [%] for 5 min,  $\alpha=95\%$ .

	Season				Season		
Method	Summer	Autumn	Winter	Method	Summer	Autumn	Winter
Optimal	34.2	16.7	24.0	Optimal	14.2	6.62	11.2
Silhouette	54.7	25.7	31.5	Silhouette	17.9	6.70	14.5
Exhaustive Search	37.9	20.7	24.0	Exhaustive Search	16.2	15.1	23.4

(a) Method A

(b) Method B

518 5.5.2. Comparison with Benchmark Methods

519 First, we present the advantage given by the k-means clustering for forecast-  
 520 ing of 1 and 5 minutes ahead. Table 8 shows the comparison with the simple  
 521 quantiles extraction for the original and differentiated time series (Quantiles A  
 522 and B, respectively) for 1 minute forecast horizon, as explained for the ultra-  
 523 short term analysis. Table 9 shows the same comparison for 5 minutes time  
 524 horizon. For Method A, we apply the exhaustive search while for B we apply  
 525 the silhouette analysis. For all the cases, we can see an improvement coming  
 526 from the k-means clustering with respect to the simple quantile extraction.  
 527 The last row shows the results obtained by implementing the ARIMA model and  
 528 assuming a Gaussian distribution of the point forecast error (ARIMA+GAUSS).  
 529 For these forecast horizons, the PICP is slightly lower than the target confidence

530 level, with values of PINAW that are however higher than those returned by  
 531 Method B.

Table 8: PICP-PINAW-CWC [%] for a time horizon of 1 minute,  $\alpha=95\%$ .

Method	Season		
	Summer	Autumn	Winter
Method A	90.1-10.2-25.2	90.8-4.81-11.4	88.8-8.87-23.4
Quantiles A	94.7-56.7-112	93.1-29.5-61.9	95.8-34.7-34.7
Method B	96.9-10.5-10.5	97.5-3.26-3.26	97.8-9.1-9.1
Quantiles B	89.7-13.8-34.8	90.6-6.1-14.6	91.7-6.73-15.3
ARIMA+GAUSS	93.4-19.2-40.3	94.0-8.13-16.5	95.6-10.6-10.6

Table 9: PICP-PINAW-CWC [%] for a time horizon of 5 minutes,  $\alpha=95\%$ .

Method	Season		
	Summer	Autumn	Winter
Method A	91.5-16.5-37.9	86.7-6.96-20.7	96.1-24.0-24.0
Quantiles A	94.7-55.6-110	93.1-28.5-60.6	95.8-33.7-33.7
Method B	96.7-17.9-17.9	96.2-6.70-6.70	96.1-14.5-14.5
Quantiles B	89.4-25.9-66.4	89.5-13.6-34.0	91.2-16.0-36.0
ARIMA+GAUSS	91.7-28.9-65.8	92.3-14.1-31.4	95.0-19.0-19.0

532 It is difficult to compare the proposed method with results available in the  
 533 literature due to the different GHI measurements (characterized by dissimilar  
 534 climatology). Similar results are obtained in [21] for 5 minutes ahead GHI  
 535 forecast, where a probability coverage of  $\approx 95\%$  and PINAW of  $\approx 8\%$ . In [21] a  
 536 dataset of 1 year is considered. The percentage of periods of high volatility (i.e.

537 with  $\Delta K$  higher than 0.5) is  $\approx 0.3 - 0.6\%$  while in our datasets is  $\approx 0.9 - 1.5\%$ .

### 538 5.6. From Ultra-short to Short Term Forecasts

539 For sub-second time horizons, the best performance is obtained by apply-  
540 ing the exhaustive search coupled with Method B. On the contrary, for higher  
541 forecast horizons, the silhouette analysis coupled with Method B is the most  
542 performing one.

543 For all the considered horizons, differentiating the time series has a positive  
544 effect on the final performance and allows to have a PICP higher than or equal  
545 to  $\alpha$ . However, the improvement coming from the differentiation decreases with  
546 increasing forecast horizons. Indeed, performance of Method A worsens less  
547 than those of Method B when increasing the forecast horizon.

548 To complete the analysis, Fig. 11 shows our metrics as a function of the  
549 forecast horizon and for different confidence levels: 85%, 95%, and 99%. We  
550 can see that PICP (left side) is always higher or equal to the target confidence  
551 level. Furthermore, the value of PINAW (right side) increases with the forecast  
552 horizon (to account for the higher uncertainty) and increases with  $\alpha$ , i.e. the  
553 method adapts the bound widths to guarantee the target coverage.

554 Fig. 12 shows the PIs and the actual realizations obtained for 500 ms, 1 and  
555 5 minutes forecast horizons, respectively. Method B is applied and the target  
556 confidence level is 99%. A day of high variability and a clear-sky day from the  
557 Winter period are selected for the comparison. The corresponding values of  
558 PINAW are shown in Table 10, for the two days respectively, showing that the  
559 PIs are narrower for the clear-sky day where the variability is lower. For the  
560 clear-sky day at 500 ms a zoom is added since the PIs and the points are not  
561 easily distinguishable. We can see that larger intervals are associated to higher  
562 time horizons, this reflecting the higher uncertainty.

563 In Fig. 13 the same days of Fig. 12 are selected and PIs are plotted for  
564 different confidence level (99%, 95% and 85%). It is interesting to notice that  
565 for the considered clear-sky day a target confidence level of 85% is enough to  
566 have all the measurements inside the PI.

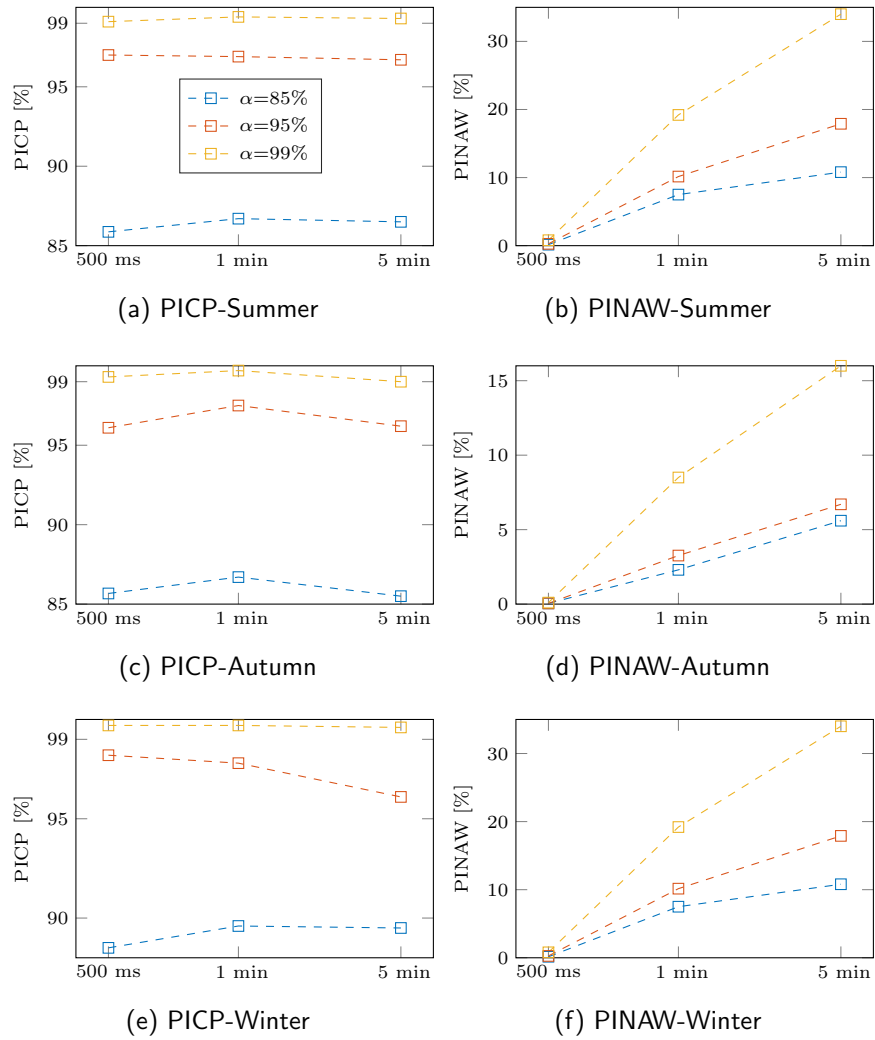


Figure 11: PICP [%] and PINAW [%] are shown for the Summer, Autumn, and Winter periods and different target confidence levels  $\alpha$ .

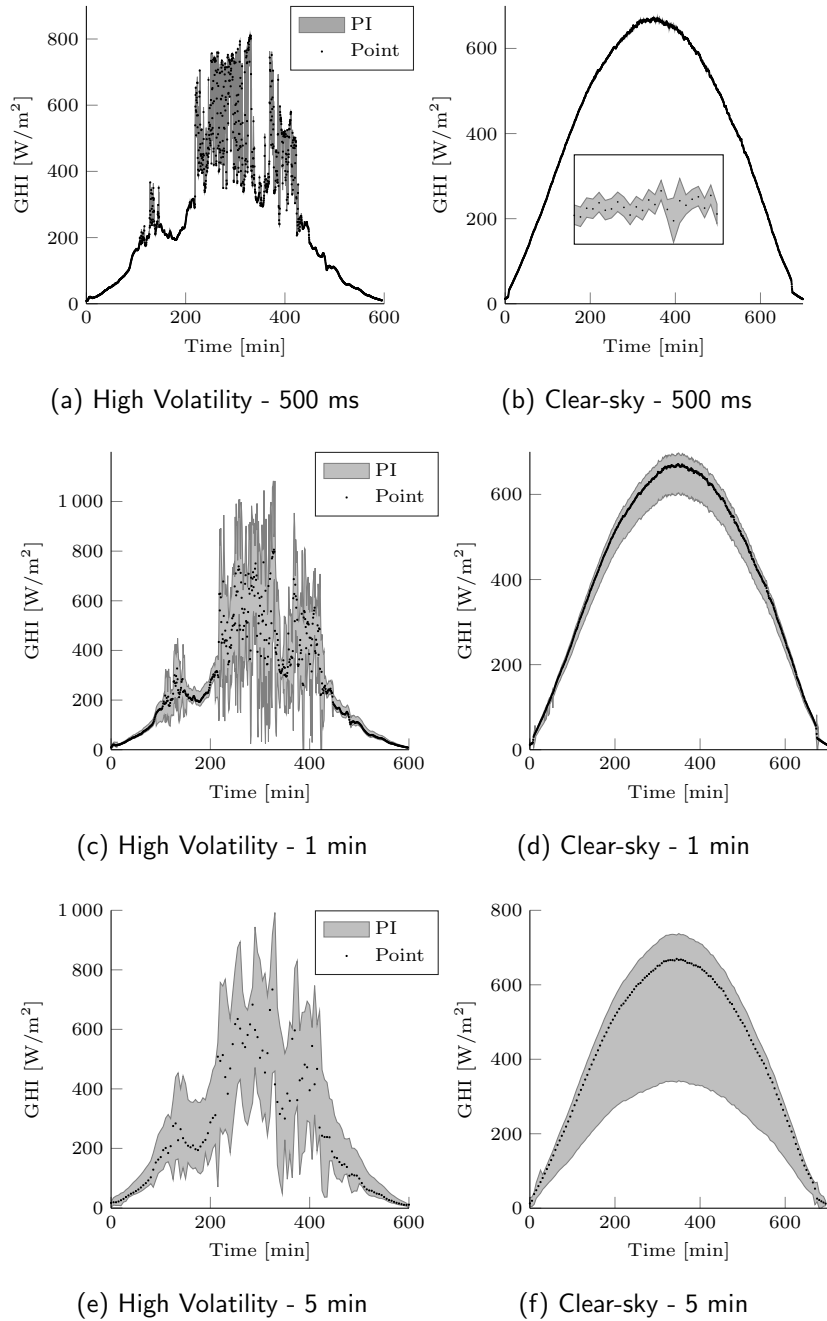


Figure 12: PIs and realizations are shown for different forecast horizons considering daylight hours,  $\alpha = 99\%$  and Method B is applied. Two days with different weather conditions are selected from the Winter period.



Table 10: PINAW [%] is shown for the two days of Fig. 12

Forecast Horizon	Day	
	High Volatility	Clear-sky
500 ms	0.22	0.18
1 min	11.8	10.8
5 min	24.9	23.8

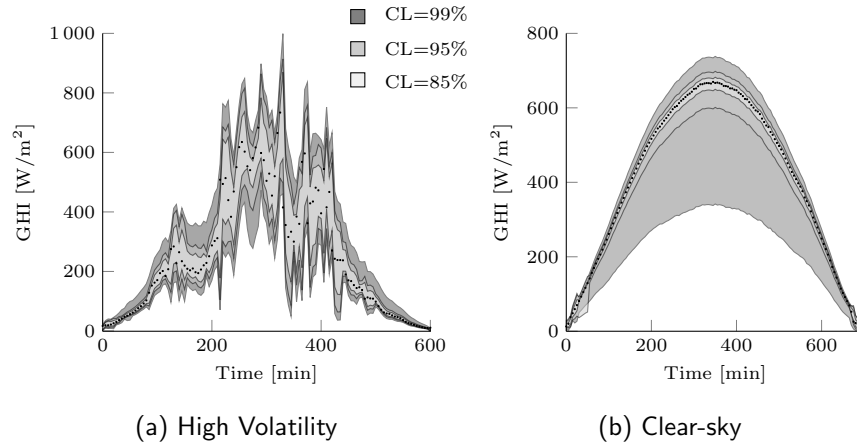


Figure 13: PIs and realizations are shown for 5 minutes time horizon and different target confidence levels, represented by different levels of shadings for the same data of Fig. 12.

567 *5.7. Reliability diagrams*

568 The objective of this last analysis is to compare the target confidence levels  
 569 with the observed ones, here represented by metric PICP.

570 The analysis is for the three considered periods and three forecast horizons  
 571 and is shown in Fig. 14. We consider Method B and the ARIMA model with  
 572 Gaussian distribution of the error and compare their performance with the ideal  
 573 behaviour, namely when the target confidence level is identical to the observed  
 574 one. As it can be seen, the confidence levels obtained for Method B exhibit an  
 575 overall good matching with the target ones, proving the capability of the method  
 576 to provide reliable predictions. In particular, Method B is always slightly over  
 577 confident with the exception of the Autumn period where is under confident for  
 578 low values of  $\alpha$ . The proposed benchmark method has lower reliability, it is over

579 confident for sub-second time horizons while its behaviour for higher horizons  
580 depends on the value of  $\alpha$ . This mismatch suggests that parametric models,  
581 with the implicit assumption of a Gaussian distribution of the error, might not  
582 be suitable.

### 583 *5.8. Execution Time Statistics*

584 Execution times are computed adopting a Matlab 2016a implementation of  
585 the algorithm on an Intel Core i7-6600U CPU 2.60GHz machine. For the anal-  
586 ysis, we consider the worst case scenario corresponding to the highest number  
587 of clusters. We select it equal to 1000 that is the maximum value of clusters  
588 returned by the analysis at 500 ms (above this value of  $k$  we do not see any  
589 performance improvement). At each time step, the overall operational time re-  
590 quired to deliver the PI is always less than 0.5 ms. The mean and the standard  
591 deviation of the computational time at each time step are 0.35 ms and 1.14 ms,  
592 respectively. The method can be used for the real-time computation of PIs at  
593 sub-second time scales, and it is expected to run even faster if developed on a  
594 dedicated industrial platform and/or in a different programming language.

## 595 **6. Conclusions**

596 The problem of quantifying the uncertainty associated with solar volatility  
597 is investigated in this work, focusing on forecast horizons that are meaningful  
598 in microgrids control applications (i.e. from sub-second up to minutes).

599 A simple method to deliver PIs for GHI is proposed and its performance  
600 assessed. The proposed technique extracts information from a limited training  
601 set: data are clustered off-line by using the well-known k-means algorithm and  
602 the quantiles of the obtained clusters are then used for PIs computation. The  
603 method does not rely on any specific point forecast technique and does not need  
604 any information from sky imaging.

605 First, a clear-sky model is implemented. It is shown that the de-trending of  
606 the time-series is advantageous for time horizons higher than the minute time-  
607 scale, when the influence of the dynamics associated to solar position becomes

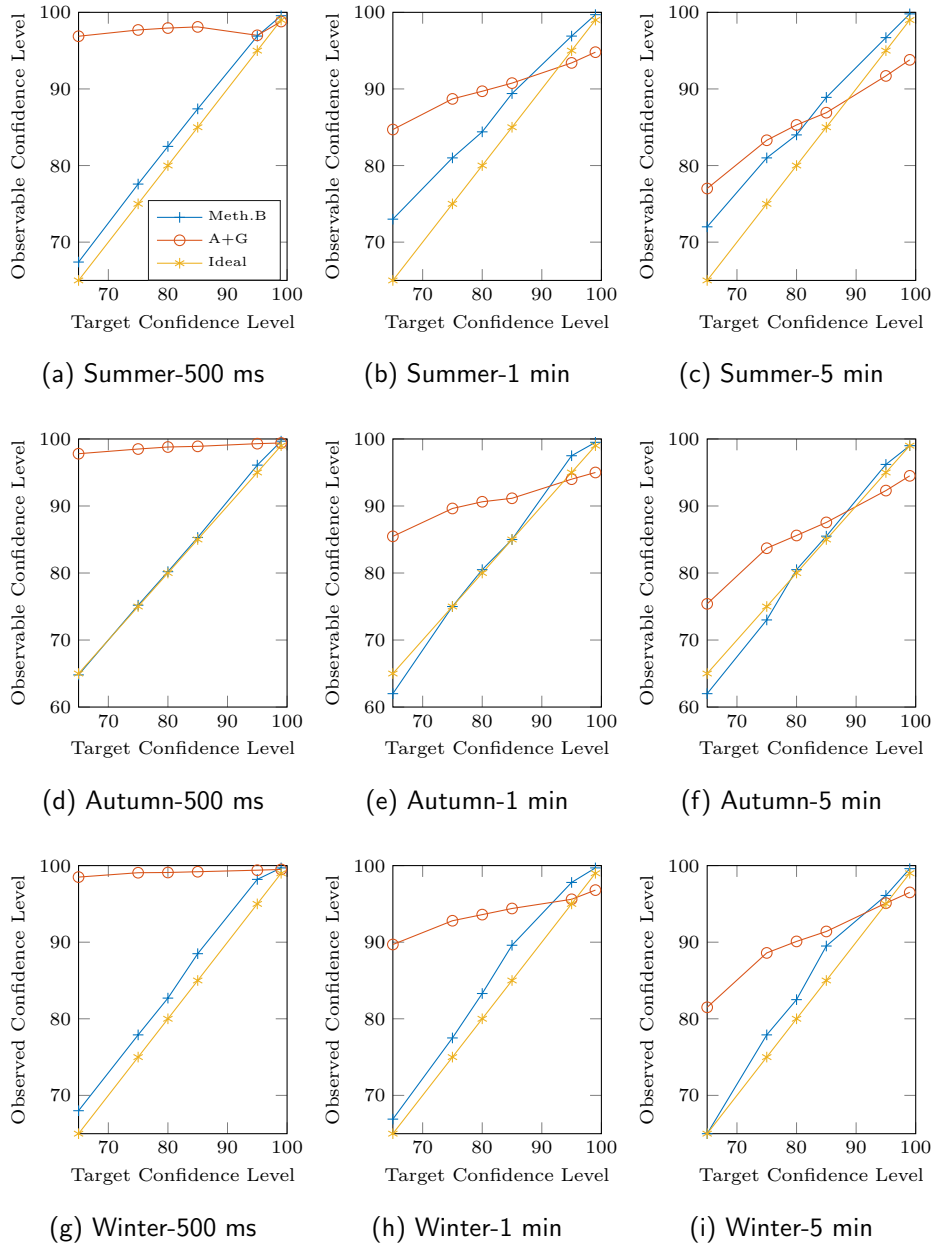


Figure 14: Reliability Diagrams for the three periods and forecast horizons.

608 non-negligible.

609 We show that the algorithm outperforms the benchmark case with simple  
610 quantiles extractions and the benchmark case considering the ARIMA model  
611 with Gaussian distribution of the point forecast error. Furthermore, perfor-  
612 mance is shown to be in line or improve those available in the literature, for all  
613 the considered forecast horizons and using a shorter and limited training set. A  
614 comparison with more sophisticated methods available in literature will be part  
615 of future work.

616 The method is applied to the original and differentiated clear-sky index time  
617 series. Results show that the benefit coming from the time series differentiation  
618 decreases while increasing the forecast horizon.

619 It is shown that the proposed method is able to adapt the widths of the PIs  
620 in order to guarantee the target coverage.

621 Thanks to its simple formulation, computational inexpensiveness and good  
622 performance at different forecast horizons, the model can be useful for providing  
623 forecast information in the field of photovoltaic generation and in the context  
624 of real-time control of microgrids.

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